

SHERLOC: Scanning Habitable Environments *with* Raman and Luminescence for Organics and Chemicals

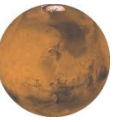
Luther Beegle, PI
Rohit Bhartia, Deputy PI
**Jet Propulsion Laboratory,
California Institute of Technology**

SHERLOC: an instrument overview.

4/23/2020

© 2016. California Institute of Technology. Government sponsorship acknowledged

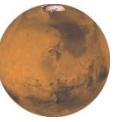




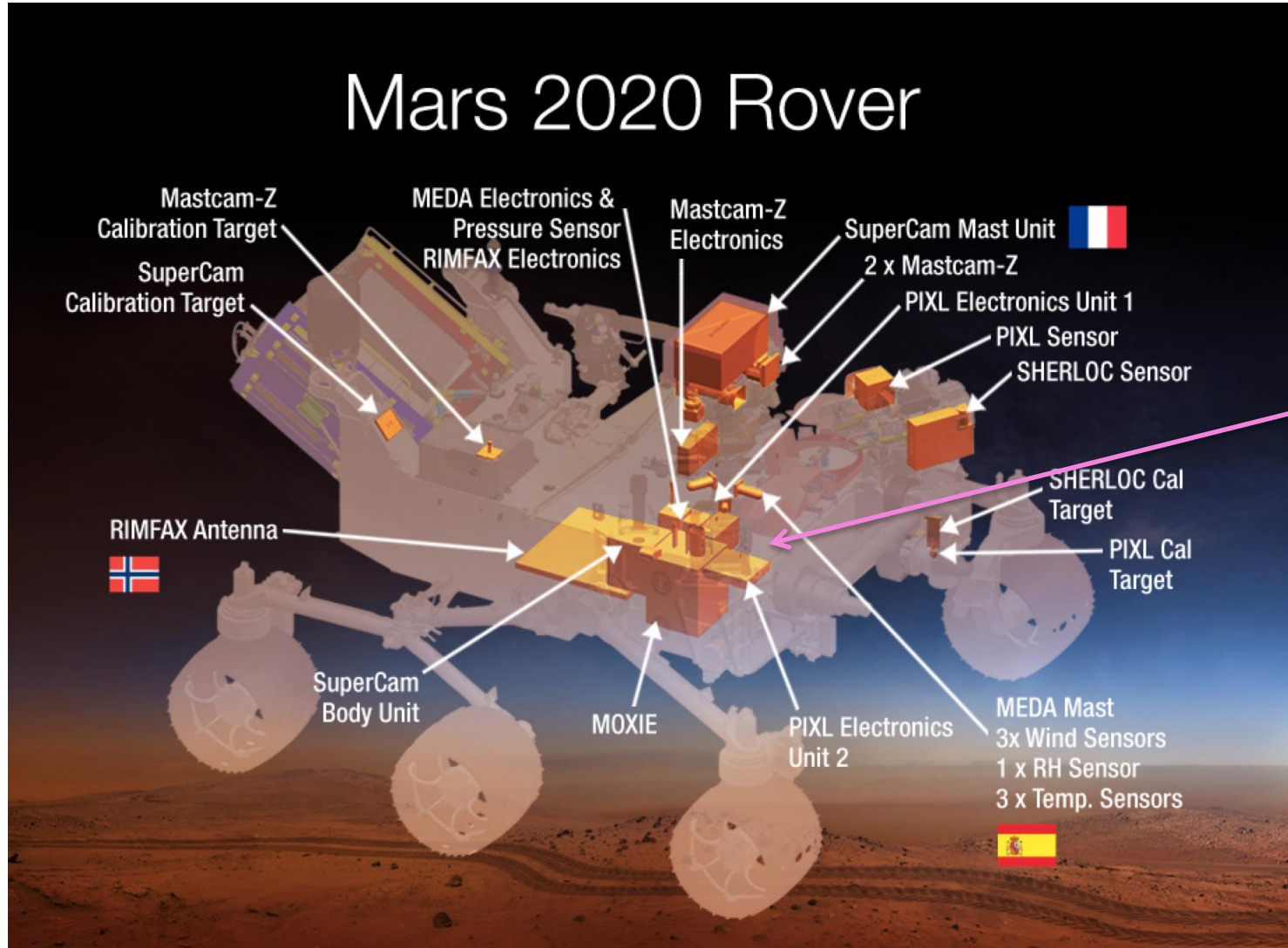
Mars 2020: First leg of potential Sample Return from Mars

- 1998: Development of the SHERLOC Concept starts
 - Funded under internal JPL R&D, NASA HQ proposals and Small Business Proposals
- Announcement of 2020 mission in December of 2012
- SDT forms January 2013
- SHERLOC design begins in January 2013
- SDT report released May 2013
 - Geology
 - Potential biosignature identification (ORGANIC identification called out)
 - Caching of well-documented samples for potential return later in the decade
 - First human-to-mars type measurements.
- Proposals Due at HQ: January 2014
- Selections July 31, 2014
- Launch in August 2020
- Landing in February 2021
- End Primary mission January 2023
- Samples potentially picked up and returned to earth 20??

The Rover

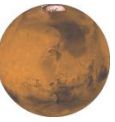


Mars 2020 Rover



**SHERLOC
Electronics
Assembly**
(not identified in
NASA press
release)

The Mars 2020 Payload



The “Spectroscopy Rover”

Mastcam – Z: High resolution zoom camera

PI: Jim Bell (ASU)

SuperCam: Elemental (LIBS), Mineralogy (IR, Visible Raman), Color imager

PI: Roger Weins (LANL)

PIXL: Elemental Maps (XRF)

PI: Abigail Allwood (JPL)

SHERLOC: *Organics and Astrobio. relevant min. (Imager+DUV Fluor/Raman)*
Microscopic to Infinity Imaging

PI: Luther Beegle (JPL)

Deputy PI: Rohit Bhartia (JPL)

MEDA: Meteorological weather station

PI: J. Manfredi (Spain)

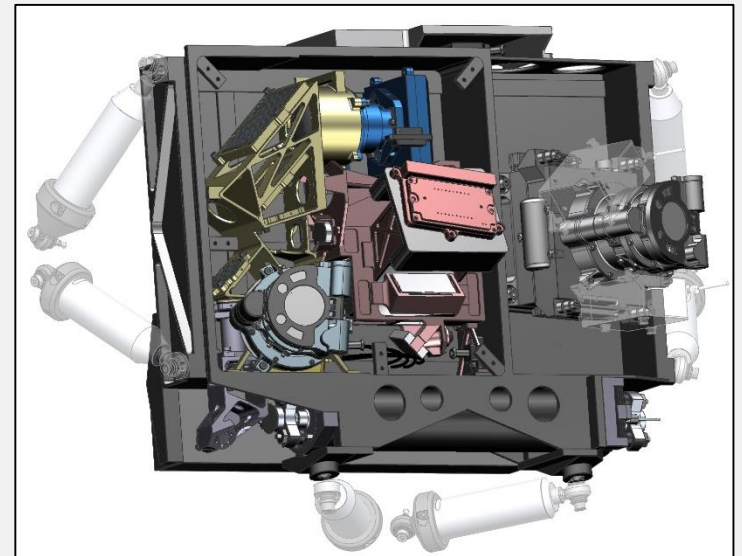
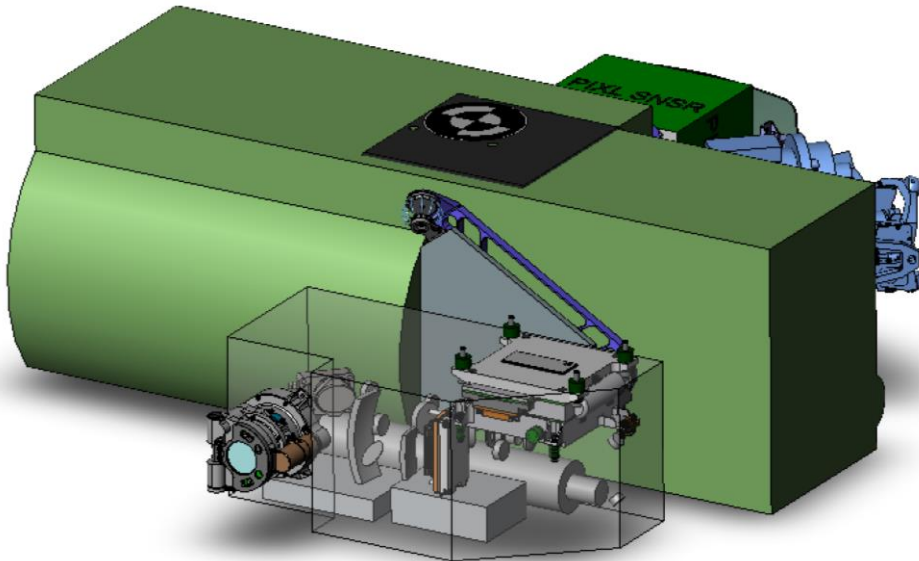
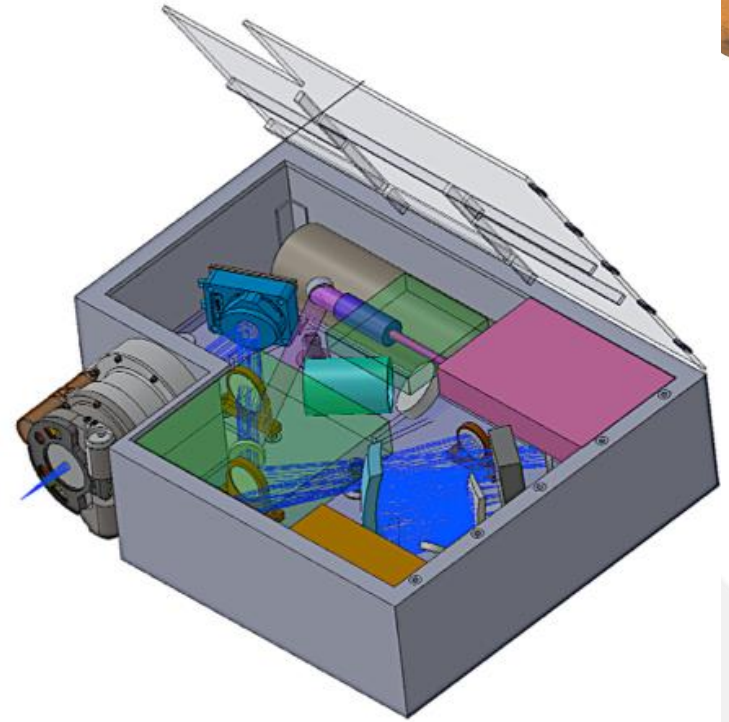
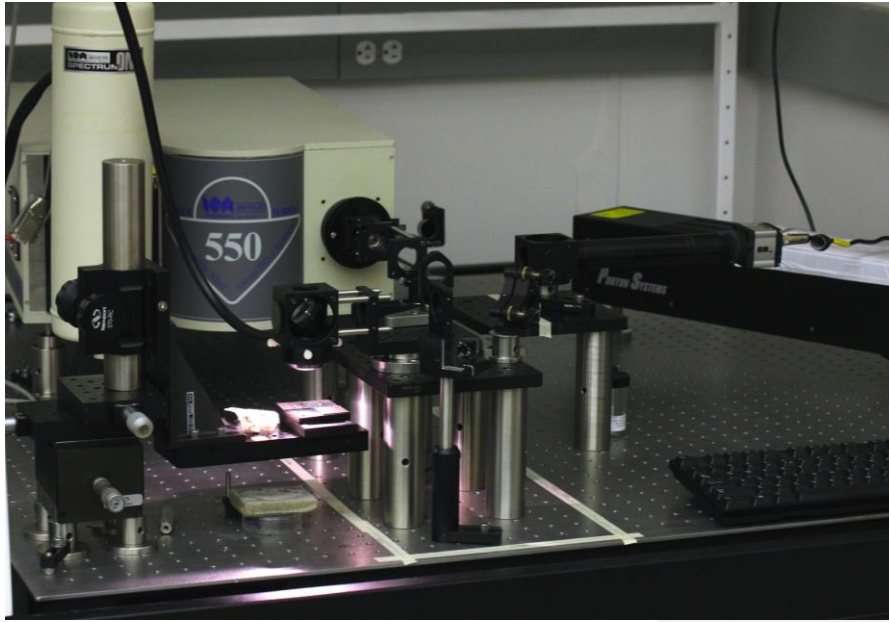
RIMFAX: Subsurface structure (GPR)

PI: Svein-Erik Hamran (Norway)

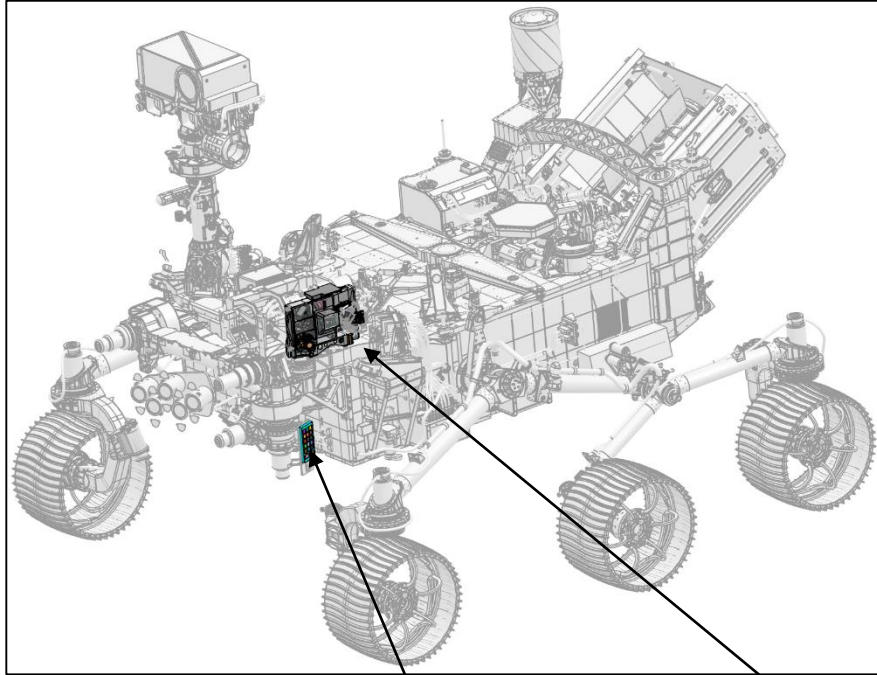
MOXIE: O₂ generation on Mars

PI: Michael Hecht (MIT)

SHERLOC



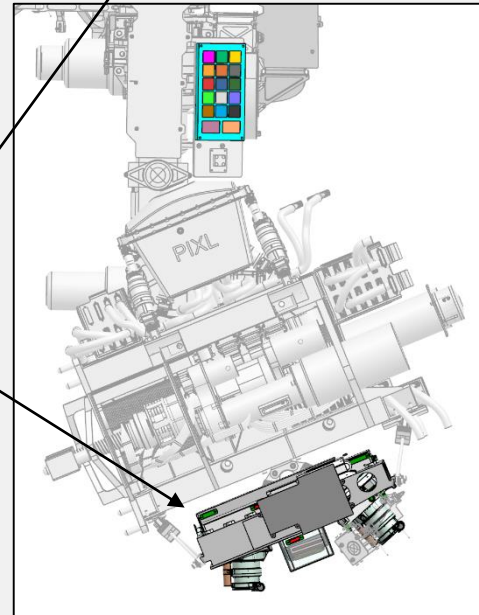
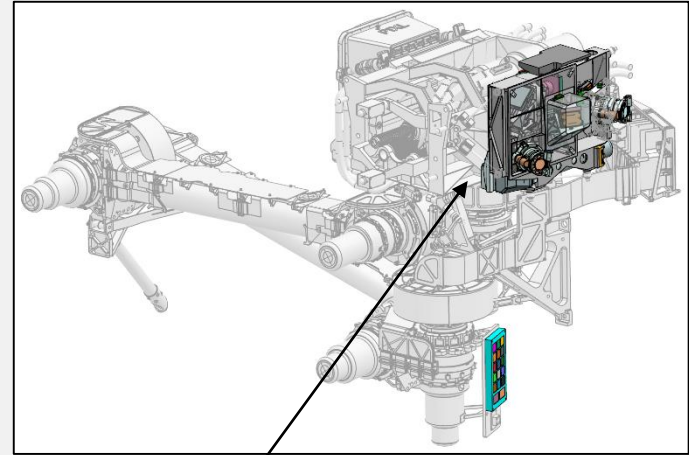
SHERLOC On the M2020 Rover



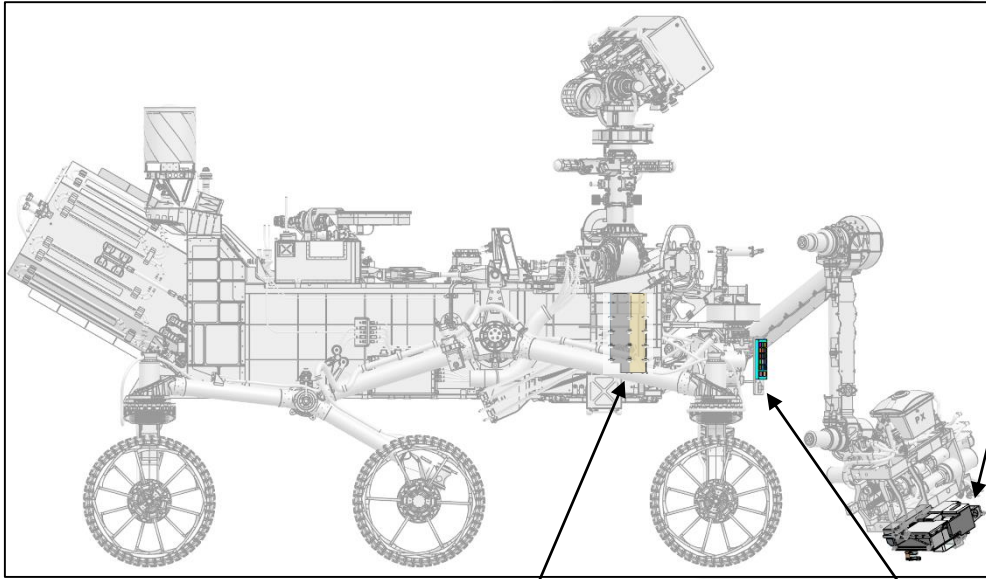
SHERLOC Body
Assembly
(Inside Rover Chassis)

SHERLOC
Calibration Target

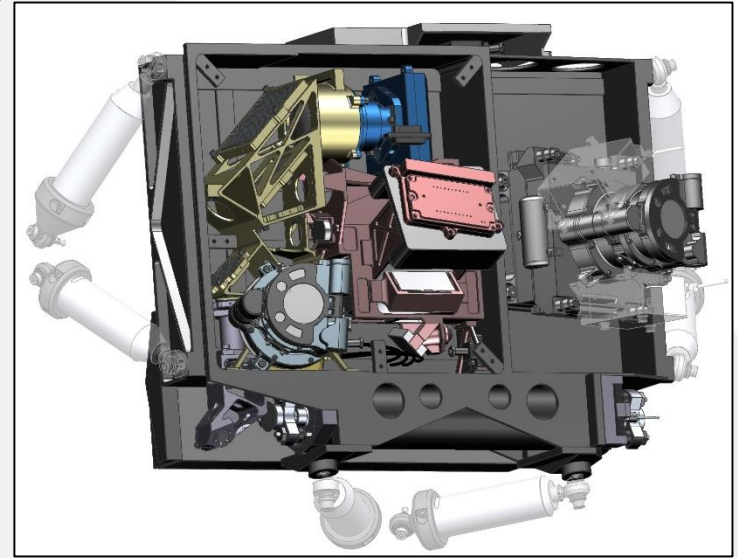
SHERLOC Turret
Assembly



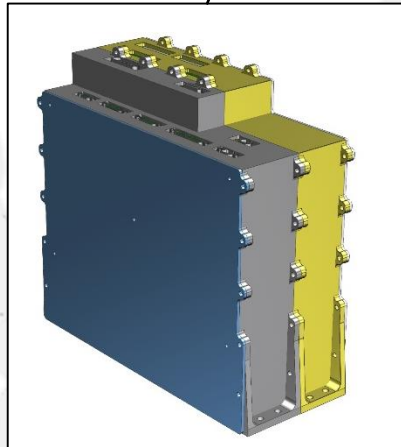
SHERLOC On the M2020 Rover



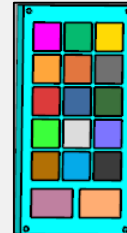
SHERLOC Turret
Assembly



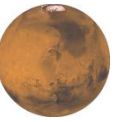
SHERLOC Body
Assembly
(Inside Rover Chassis)



SHERLOC
Calibration Target



SHERLOC Investigation Team



Astrobiology



L. Beegle
(PI)



R. Bhartia
(D-PI)



K. Williford
(JPL)



K. Nealson
(USC)



P. Conrad
(GSFC)



A. Steele
(CIW)

Laser Raman Spectroscopy



S. Asher
(Pitt Univ.)



P. Sorbon
(SETI)



M. Fries
(JSC)

Meteoritics



R. Wiens
(LANL)



A. Burton
(JSC)

Geology/Mineralogy



K. Edgett
(MSSS)



B. Ehlmann
(Caltech)



L. Kah
(Univ. Tenn)



J. Popp
(Jena)



S. Clegg
(LANL)



W. Hug
(Photonsystems)



F. Langerhorst
(Jena)

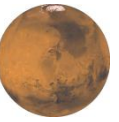


M. Minitti
(PSI)



A. Yingst
(PSI)

SHERLOC Investigation Team



Astrobiology



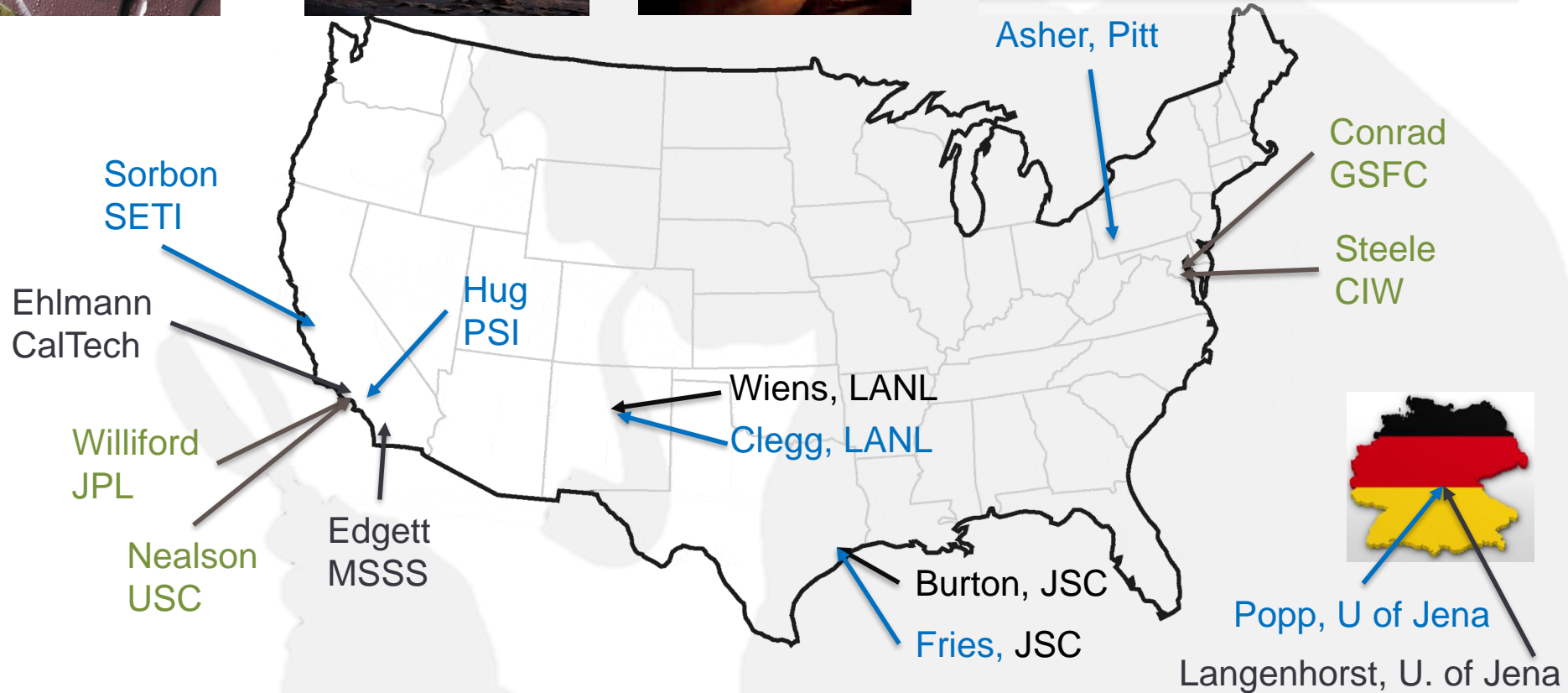
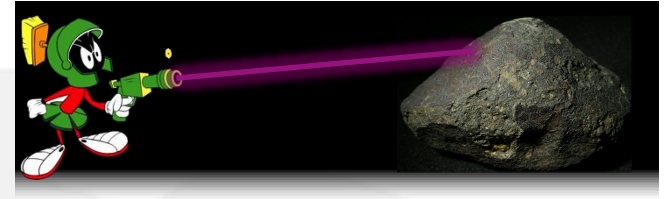
Geology



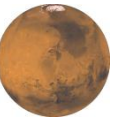
Meteoritics



Laser Raman Spectroscopy



SHERLOC Science goals

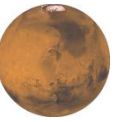


- *Assess the habitability potential of a sample and its aqueous history.*
- *Assess the availability of key elements and energy source for life (C,H,N,O,P,S etc)*
- *Determine if there are potential biosignatures preserved in Martian rocks and outcrops.*
- *Provide organic and mineral analysis for selective sample caching.*

To do this SHERLOC does the following:

- Detects and classifies organics and astrobiologically relevant minerals on the surface and near subsurface of Mars.
- Organic sensitivity of 10^{-5} to 10^{-6} w/w over the observed area
- Organic sensitivity of 10^{-2} to 10^{-4} w/w spatially resolved at $<100\mu\text{m}$
- Astrobiologically Relevant mineral detection and classification to $<100\mu\text{m}$

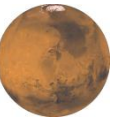
SHERLOC Science goals



- *Assess the habitability potential of a sample and its aqueous history.*
- *Assess the availability of key elements and energy source for life (C,H,N,O,P,S etc)*
- ***Determine if there are potential biosignatures preserved in Martian rocks and outcrops.***
- *Provide organic and mineral analysis for selective sample caching.*

To do this SHERLOC does the following:

- Detects and classifies organics and astrobiologically relevant minerals on the surface and near subsurface of Mars.
- Organic sensitivity of 10^{-5} to 10^{-6} w/w over the observed area
- Organic sensitivity of 10^{-2} to 10^{-4} w/w spatially resolved at $<100\mu\text{m}$
- Astrobiologically Relevant mineral detection and classification to $<100\mu\text{m}$

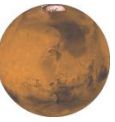


NASA's “working” Definition of Life

- A self sustaining system capable of Darwinian evolution.

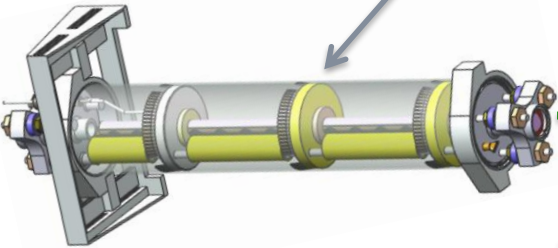


Active spectroscopy



Bringing your own source

Laser



Sample

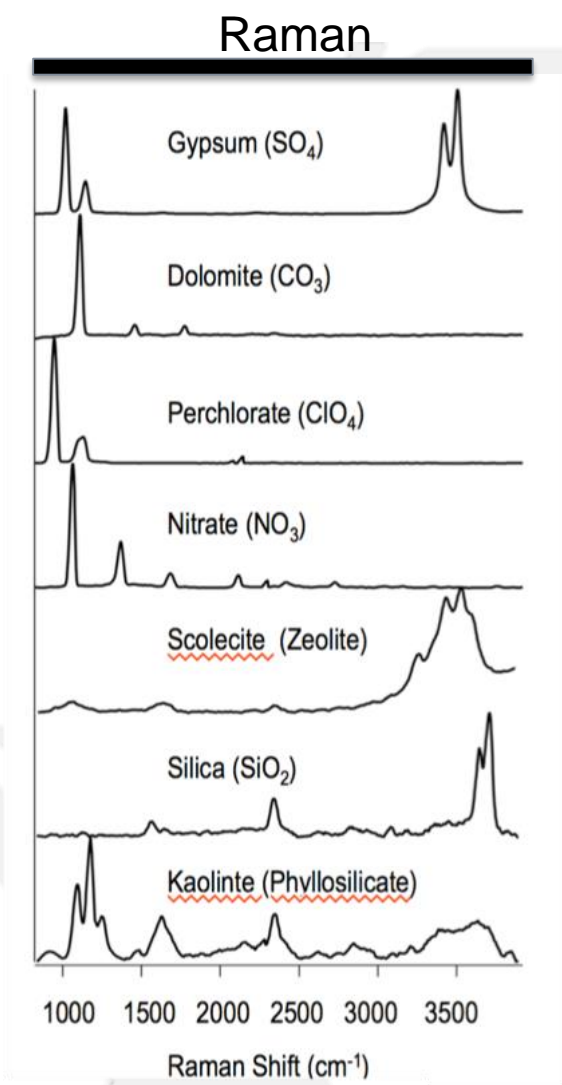


Advantages:

- Choose a wavelength of range of wavelengths to interact with a surface
- Observe interaction with the surface that are otherwise difficult to see
 - Fluorescence ← Organic rings
 - Raman ← Molecular structure
 - LIBS ← Elements
- Provides unique information about the sample

SHERLOC Capabilities

Astrobiology relevant minerals



Organics

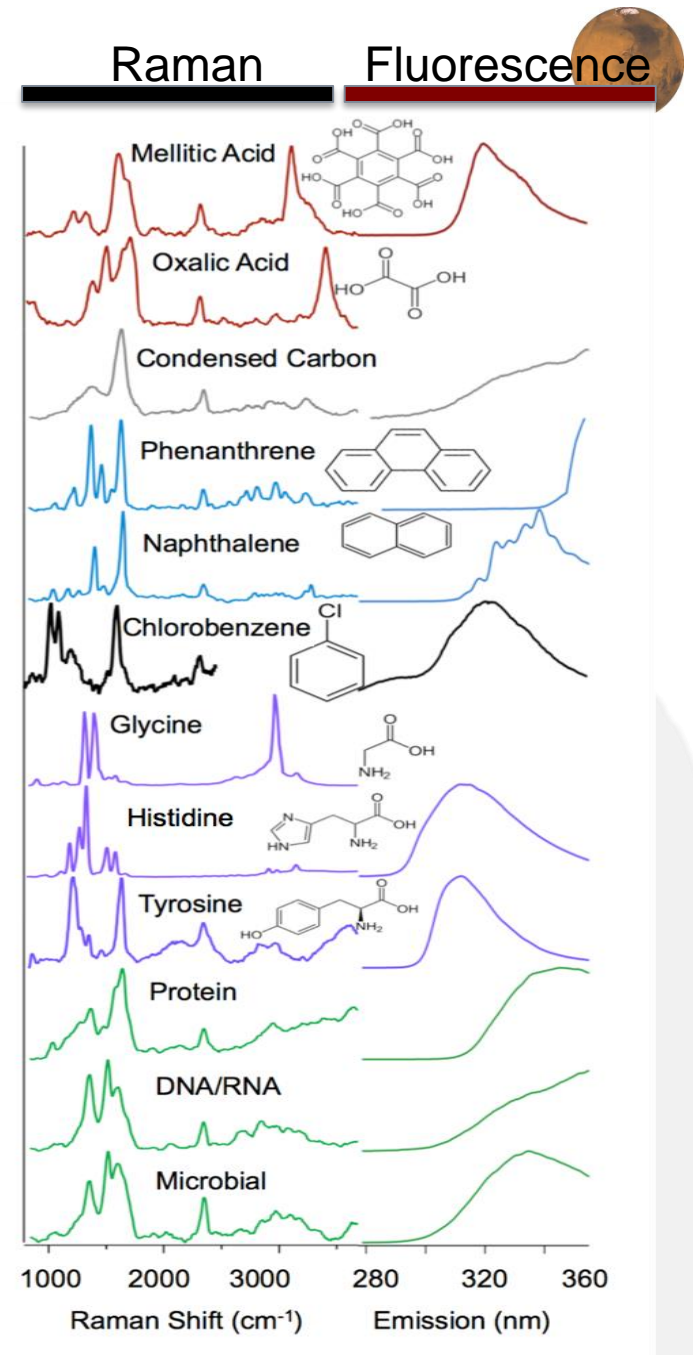
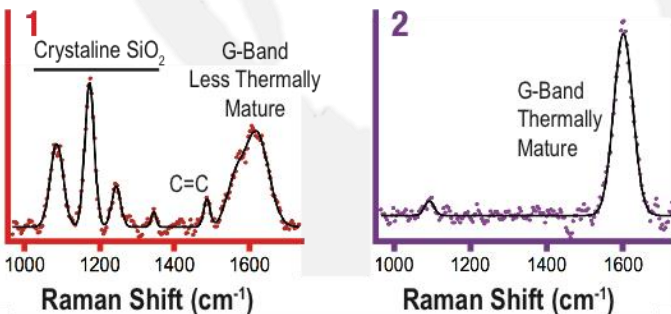
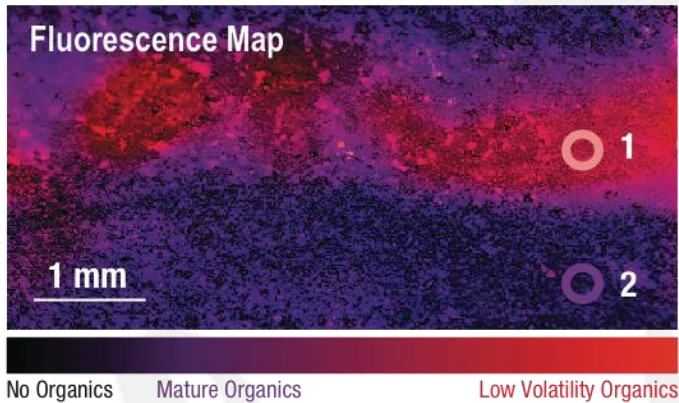
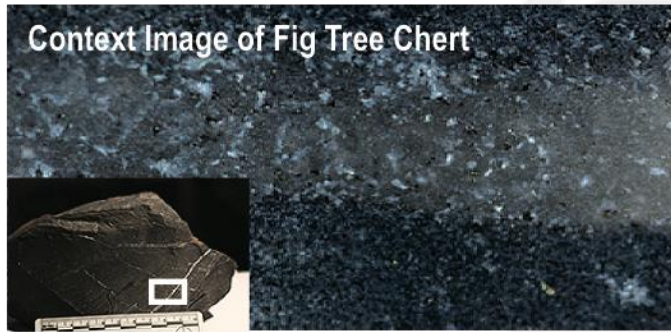
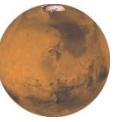


Fig Tree Chert Sample



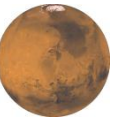
SHERLOC DEMONSTRATION: FIG TREE CHERT.

A 3.2-3.5 Ga silica-rich, low-carbon rock that has yielded microfossils and stromatolites (c.f. Lowe and Byerly 1999).

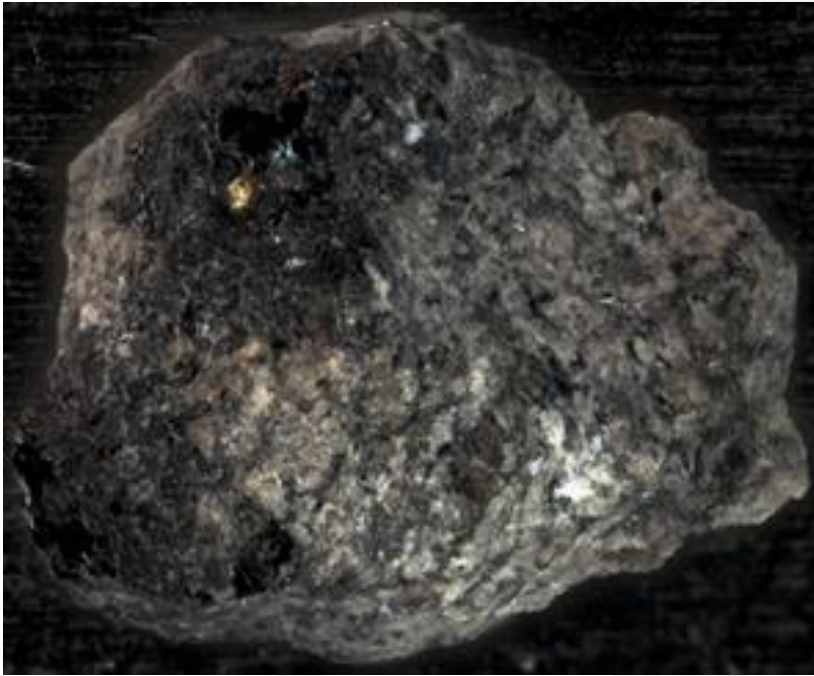
Using the SHLEORC prototype, a low-carbon standard from the Fig Tree formation was analyzed and shows both organics and mineral features that describe its environmental and aqueous history:

- 1. Crystalline SiO_2 and DUV Raman C=C and G-band positions/widths show veins are aqueous, low temperature intrusion of silica and less mature organics.*
- 2. DUV Raman G-band position/width shows the chert was exposed to temperatures $>600^\circ\text{C}$.*

Tissint Meteorite: Imaging



Context Image (30 μ m resolution)

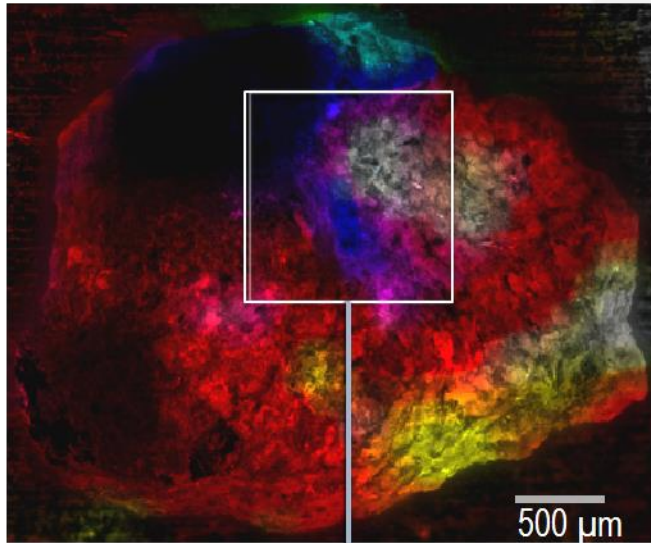


Surface Topographic Analysis

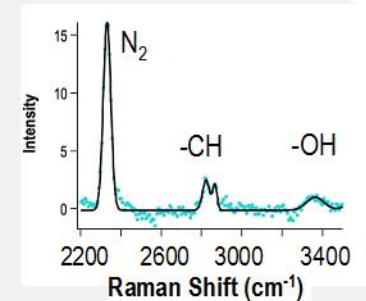
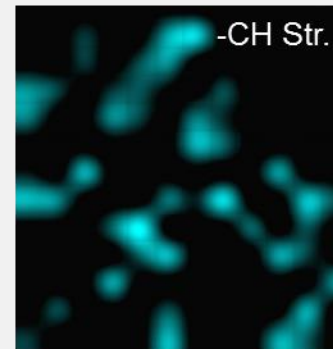
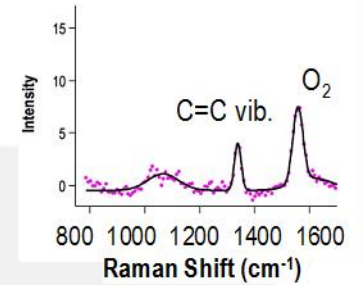
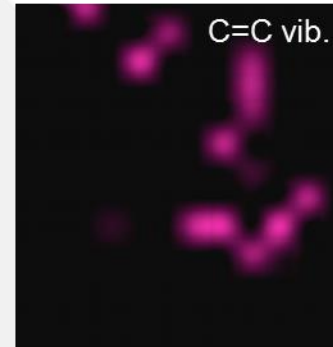


Tissint Meteorite DUV Raman Analysis

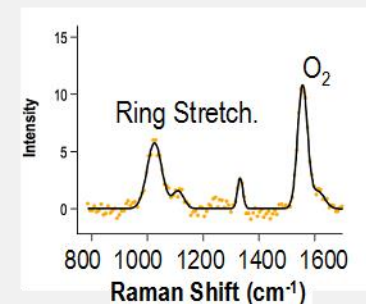
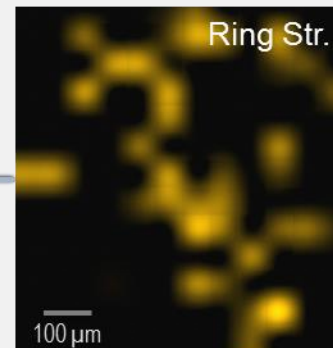
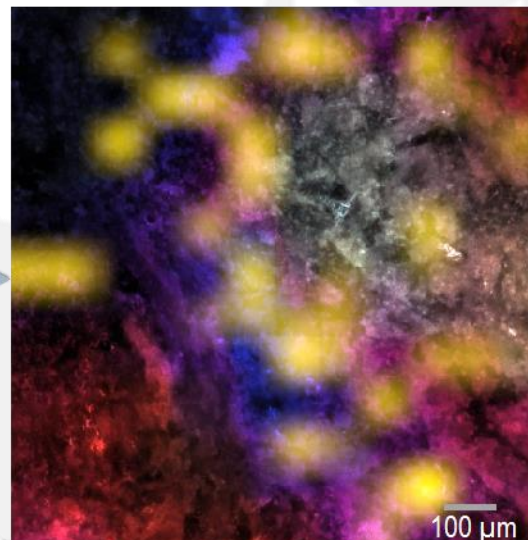
Fluorescence Map



Deep UV Raman Map

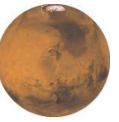


Fluorescence/Raman Map

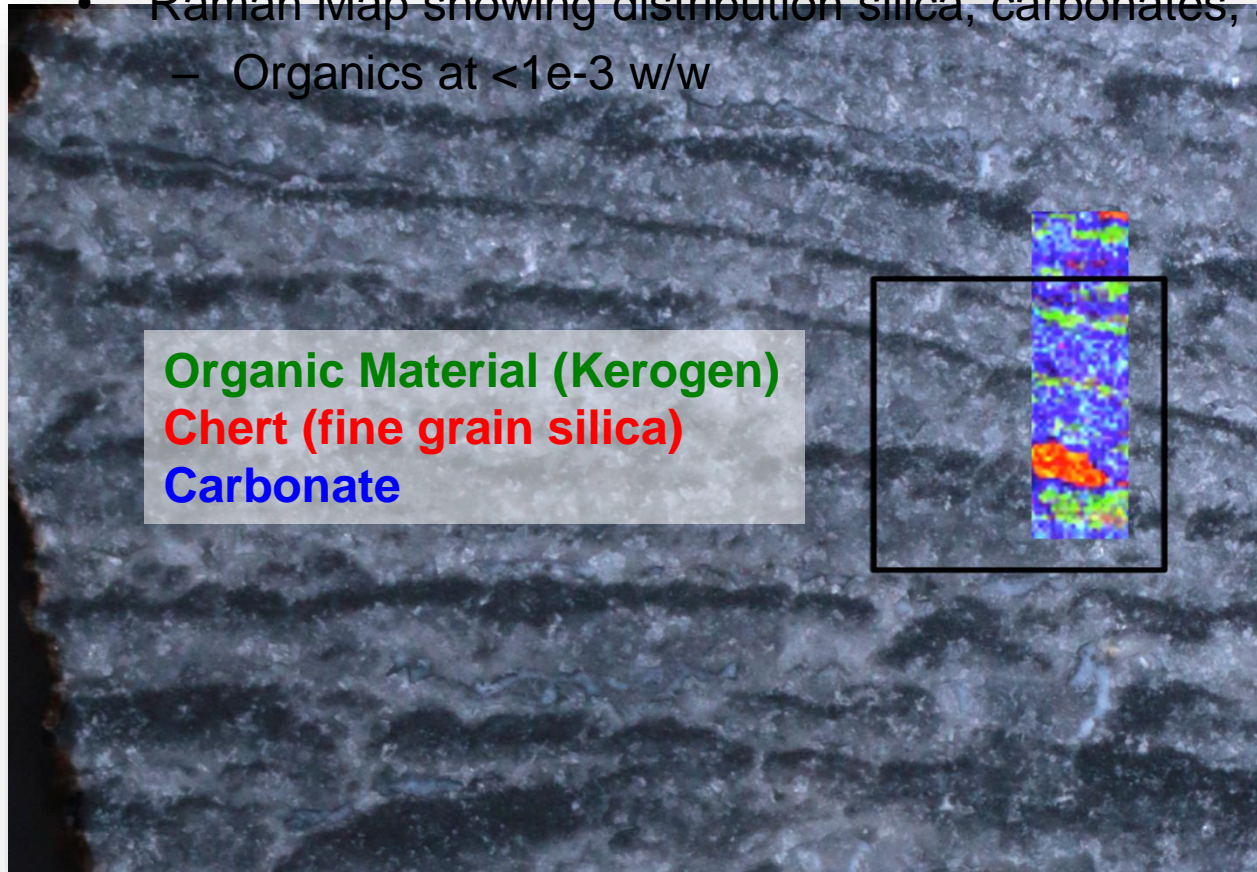


SHERLOC Science

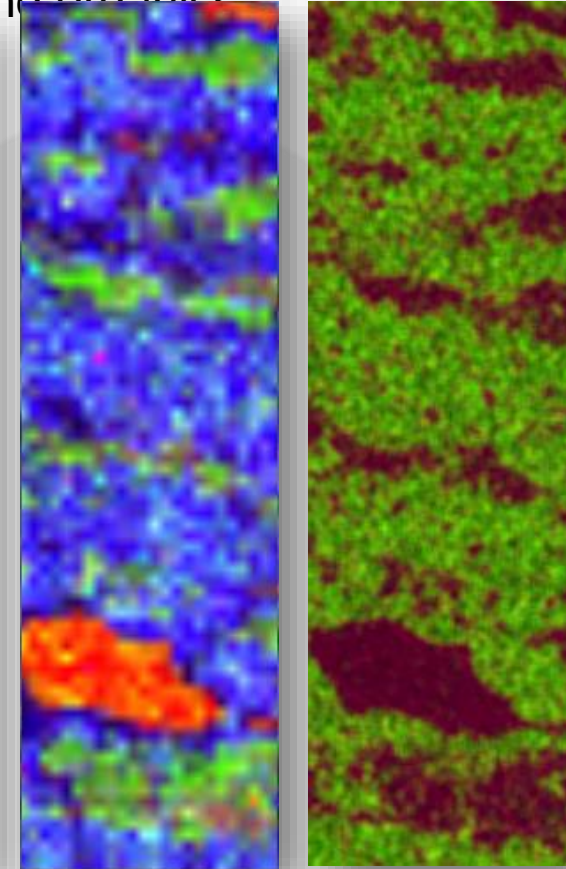
Detection and Classification of Organics and Minerals



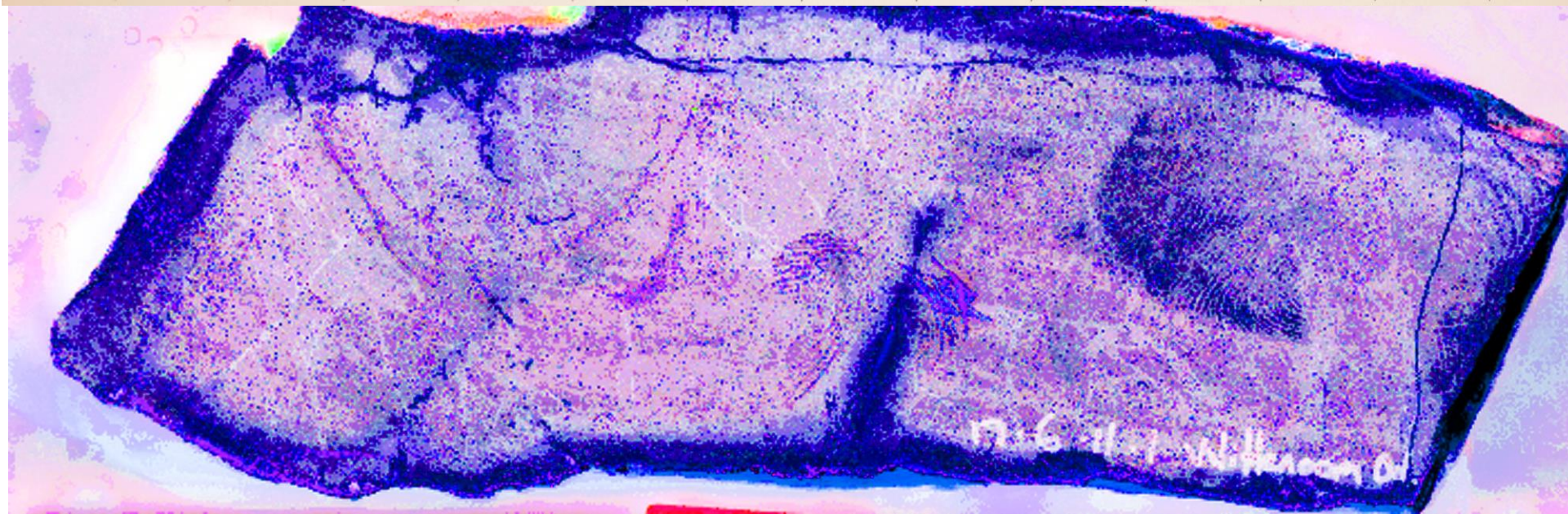
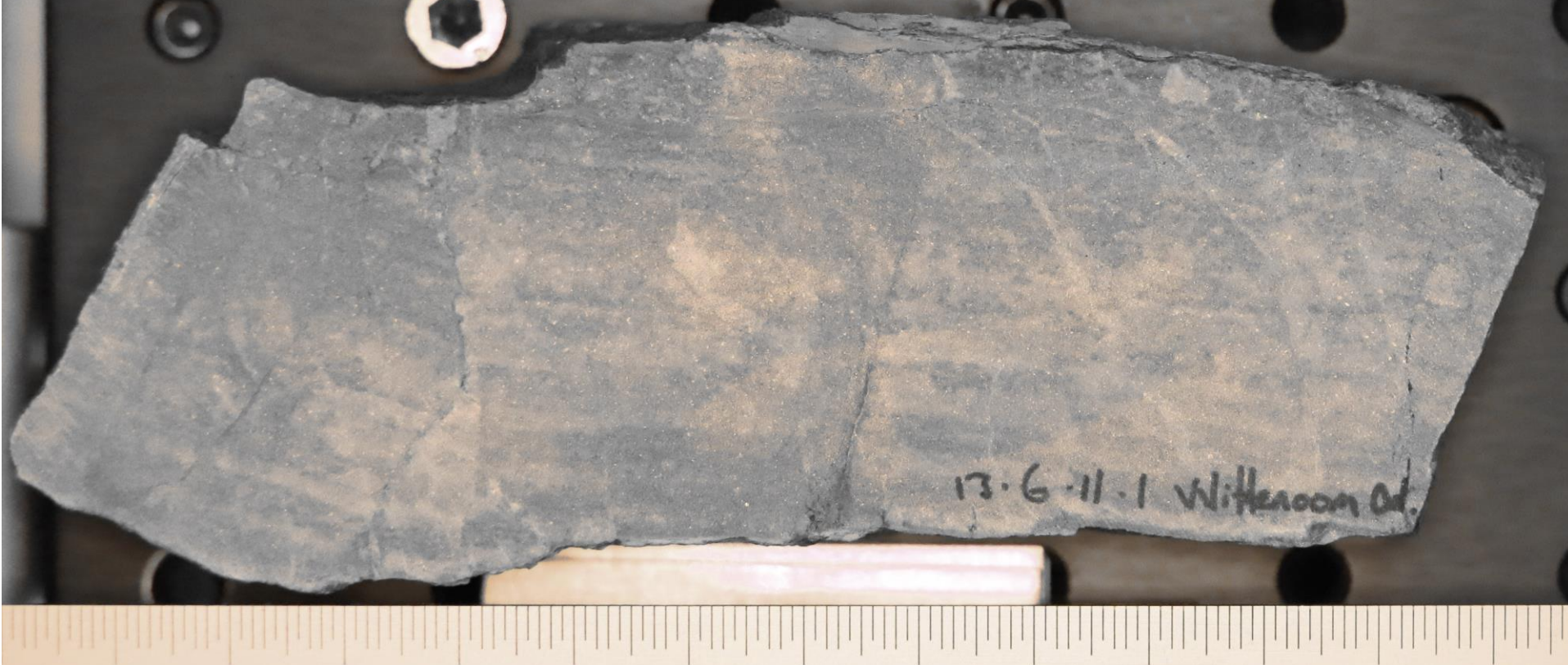
- Raman Map showing distribution silica, carbonates, and organics
 - Organics at $<1\text{e-}3$ w/w



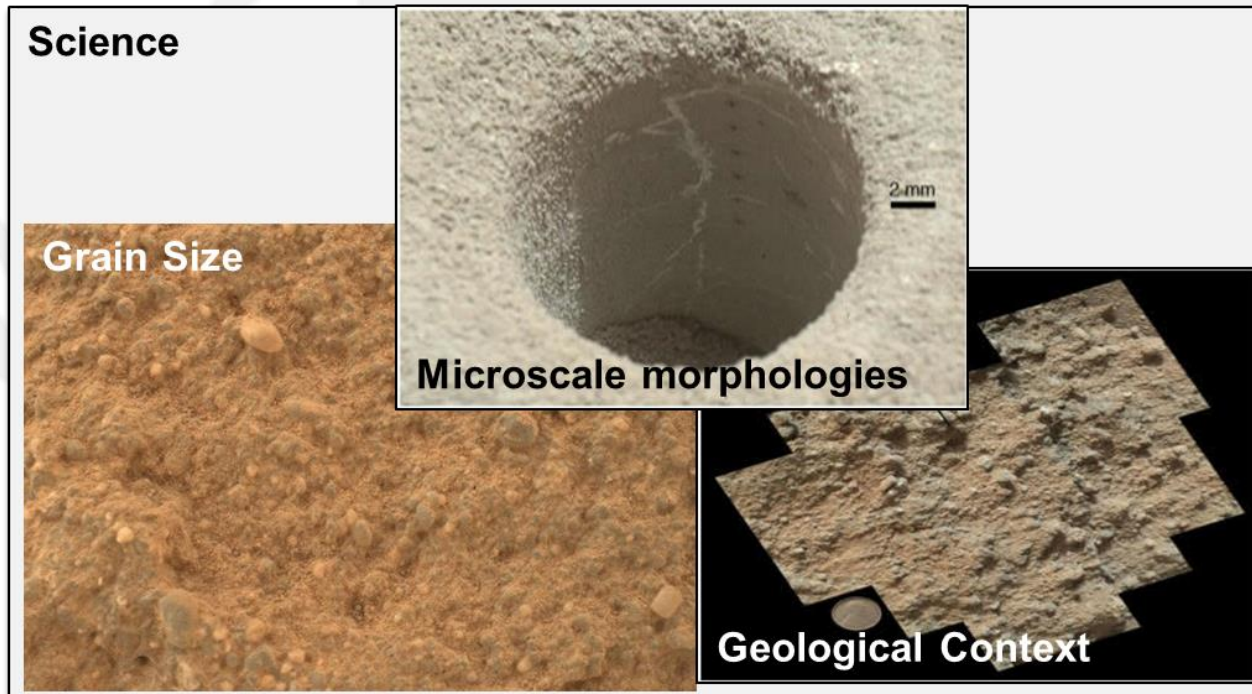
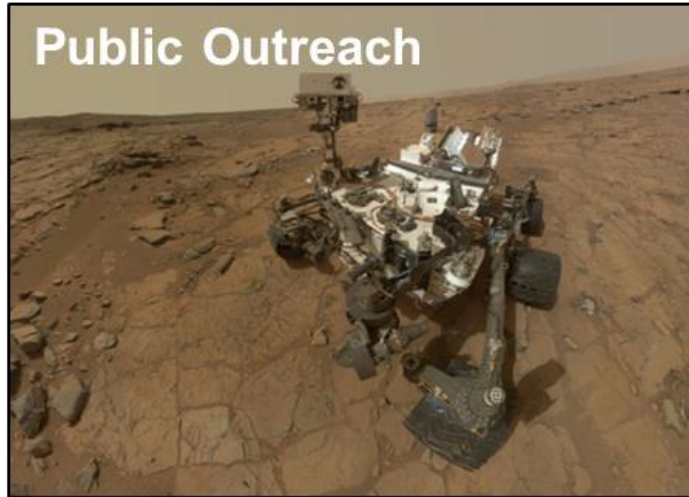
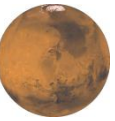
PIXL-like Data



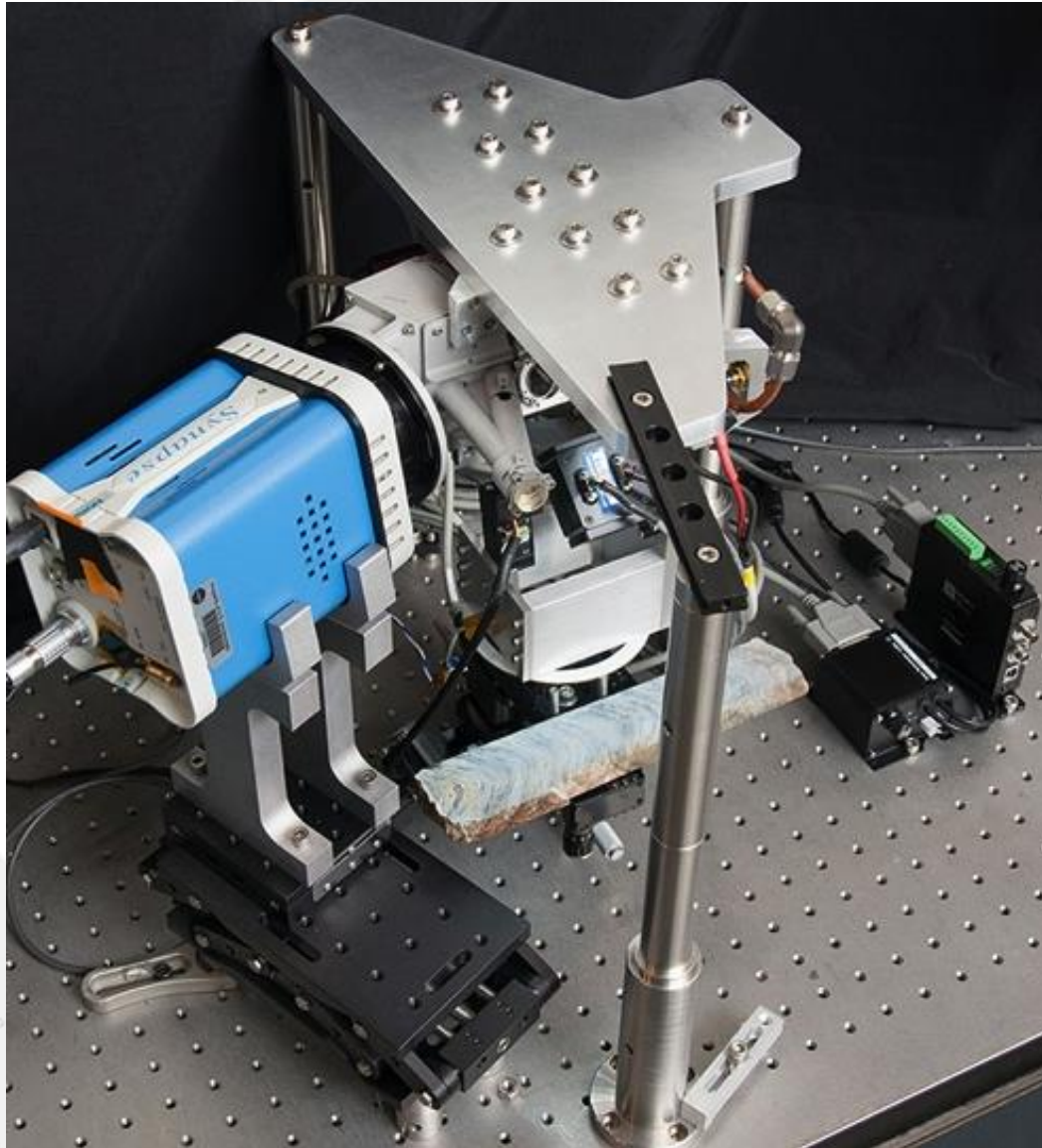
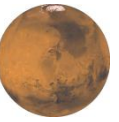
Data collected on PT1 with SHERLOC relevant parameters



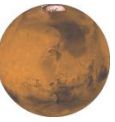
Imaging with WATSON!



Laboratory Model

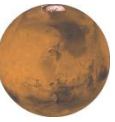


Planetary Protection & Contamination Control



- We clean the instrument
- Then we clean the room that holds the instrument
- Then we clean the instrument
- Then we clean the rover
- Then we clean the instrument
- Then we clean the rover
- Then we clean the instrument
- Then we clean the instrument
- Then we make sure the instrument is clean until we launch!

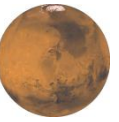
Preparing for future Humans: Space Suite material!



- HEOMD EVA Suit Targets
 - Orthofabric (outer layer of EVA suit outergarment)
 - 3oz. Teflon fabric (outer layer of glove, back of hand, gauntlet)
 - Polycarbonate (helmet visor)
 - RTV silicone (glove palm)
 - Vectran (glove palm)
 - Spectra (EVA suit structural element)
 - 6oz polyester (EVA suit restraints)



NOW TIME FOR QUESTIONS!



*4 years and 10 months to
we land on February 18th
2021*

See you at the 2021 celebration
with the first results!